Topographic Mapping of Mars:

Toward the Human Scale

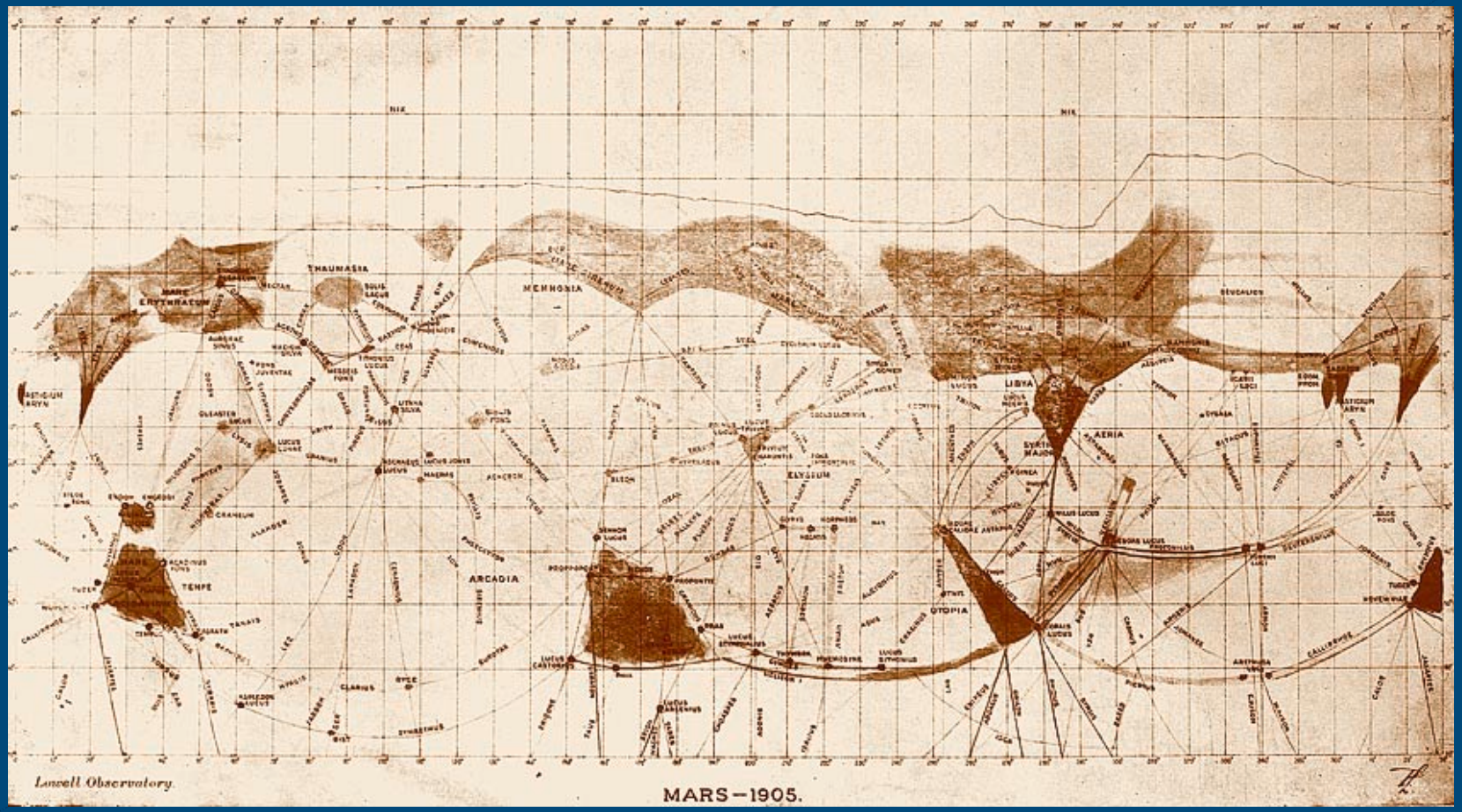
Randolph Kirk, Elpitha Howington-Kraus
Laurence Soderblom, Brent Archinal

AGU Fall Meeting
10 December 2002

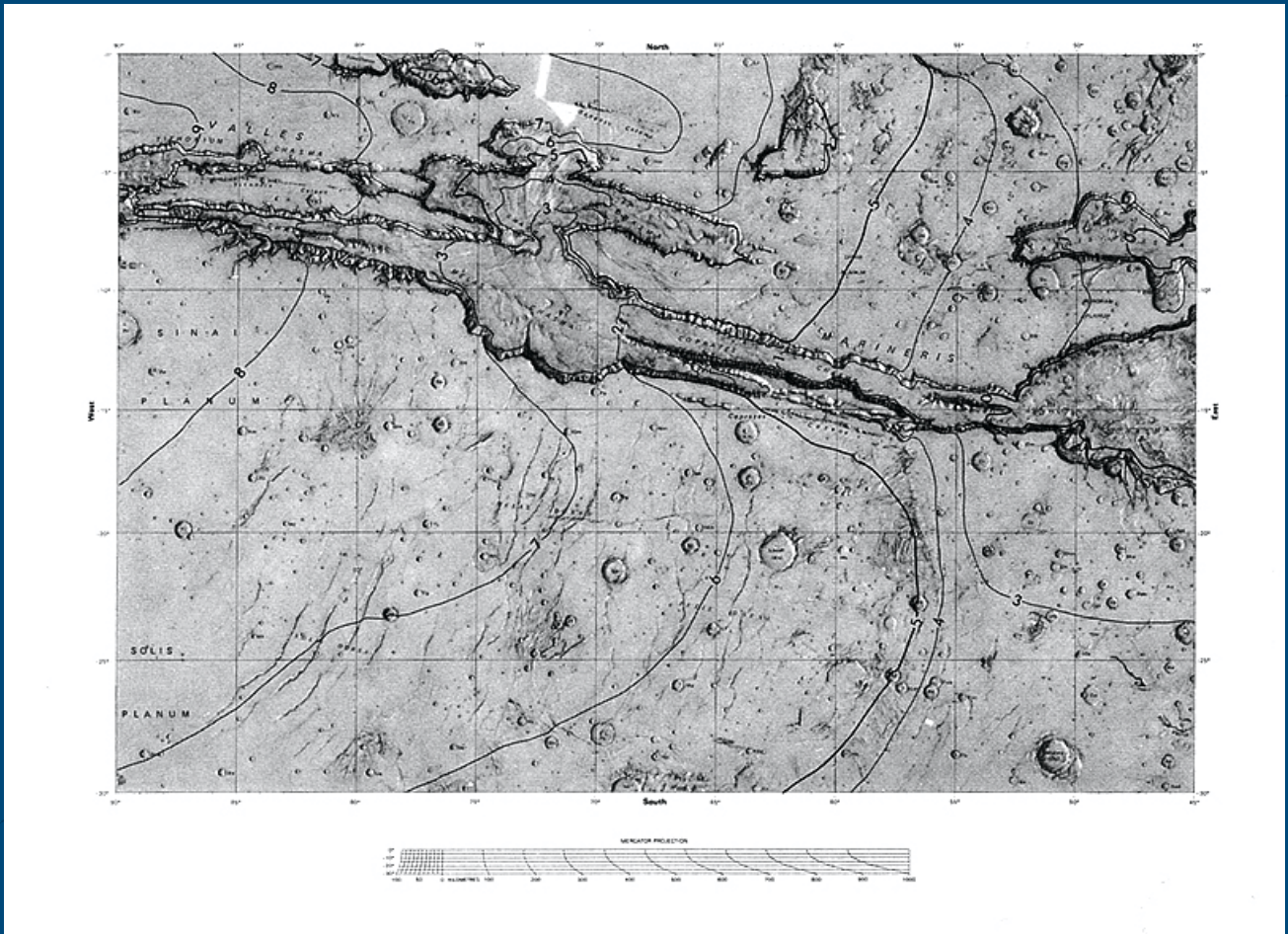
Outline

- Historical Background
- Stereomapping with MOC
- Photoclinometry
- Things to come

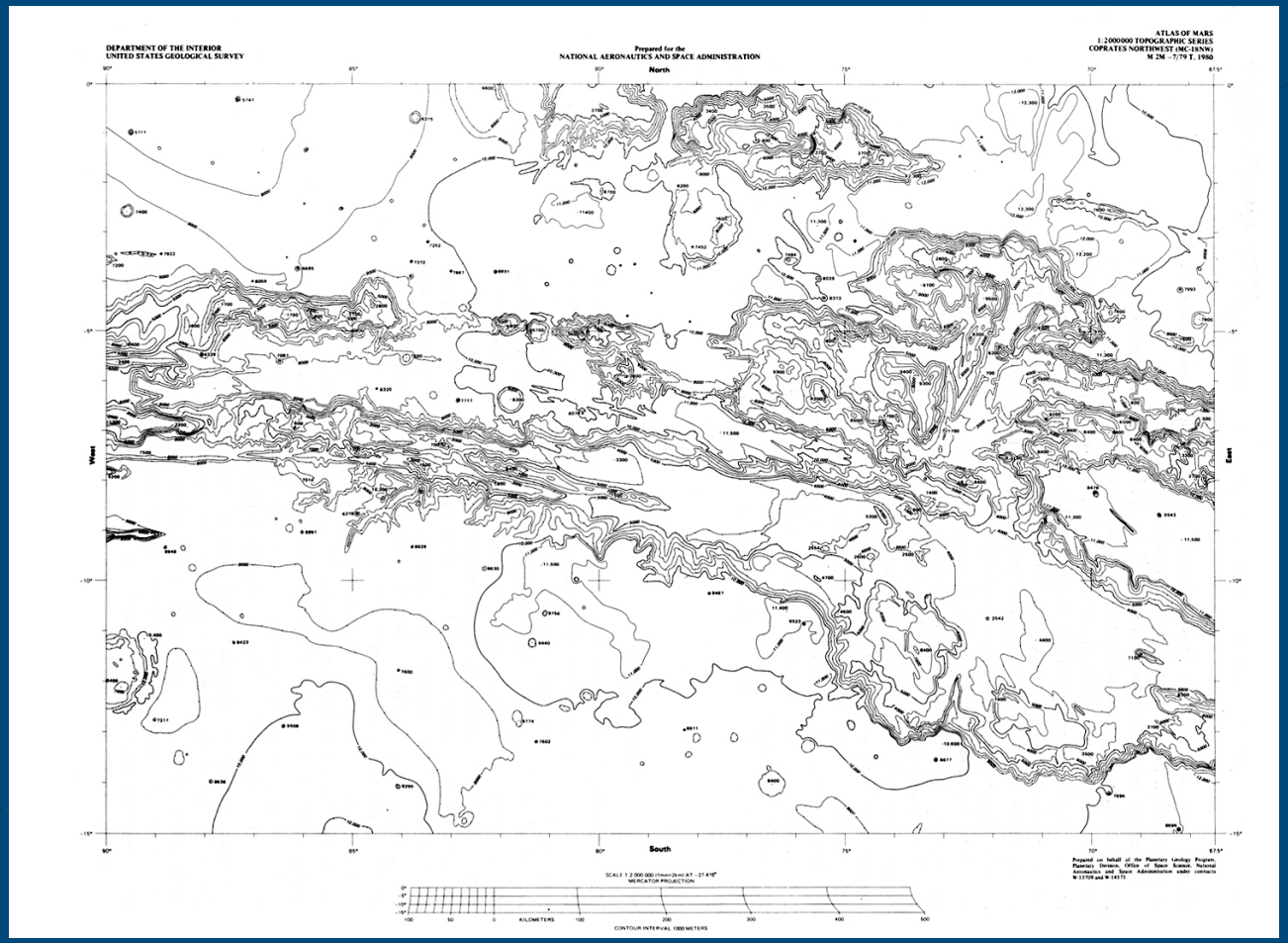
100 Years Ago



Mariner 9 Topography



Viking Stereotopographic Maps



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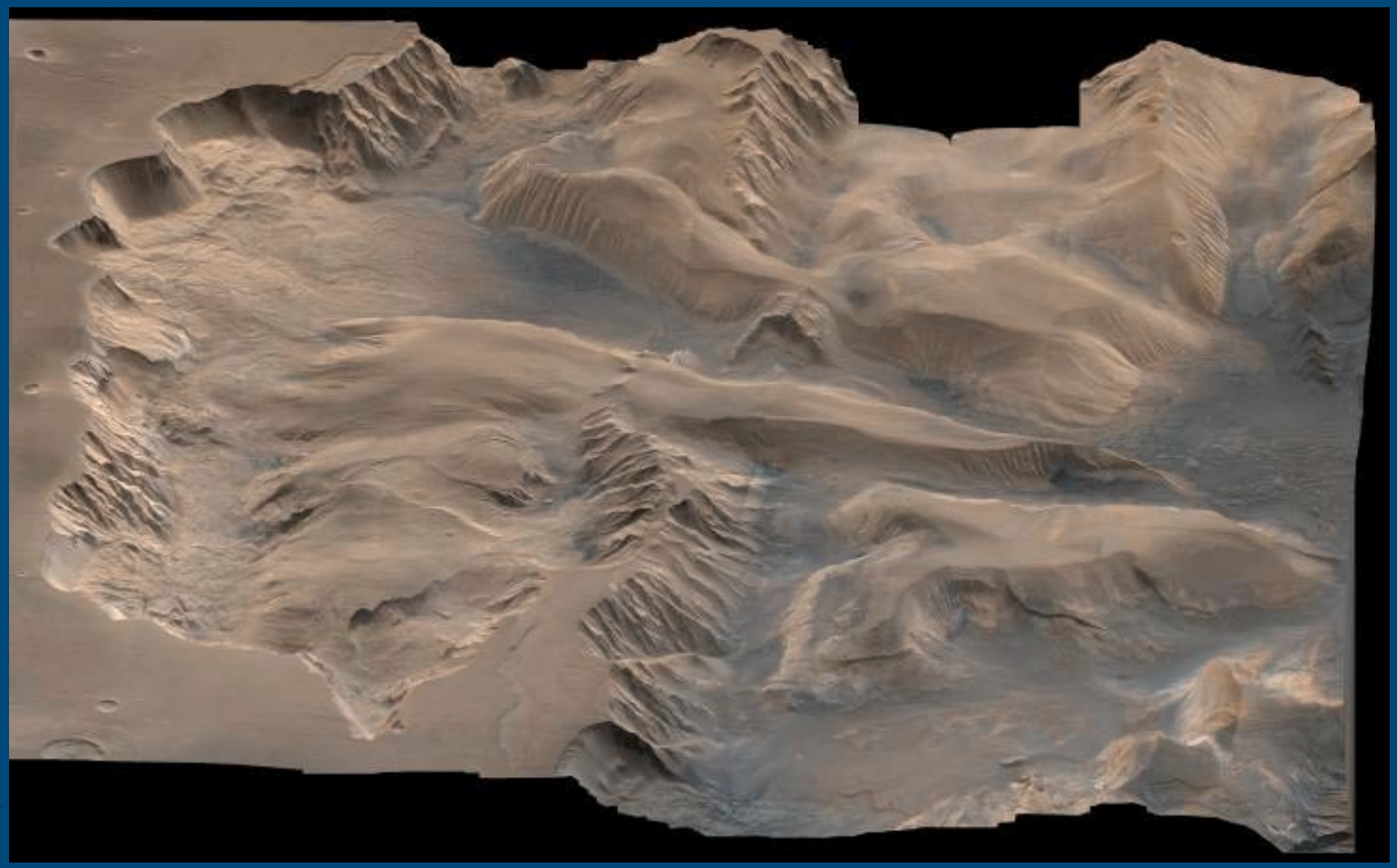

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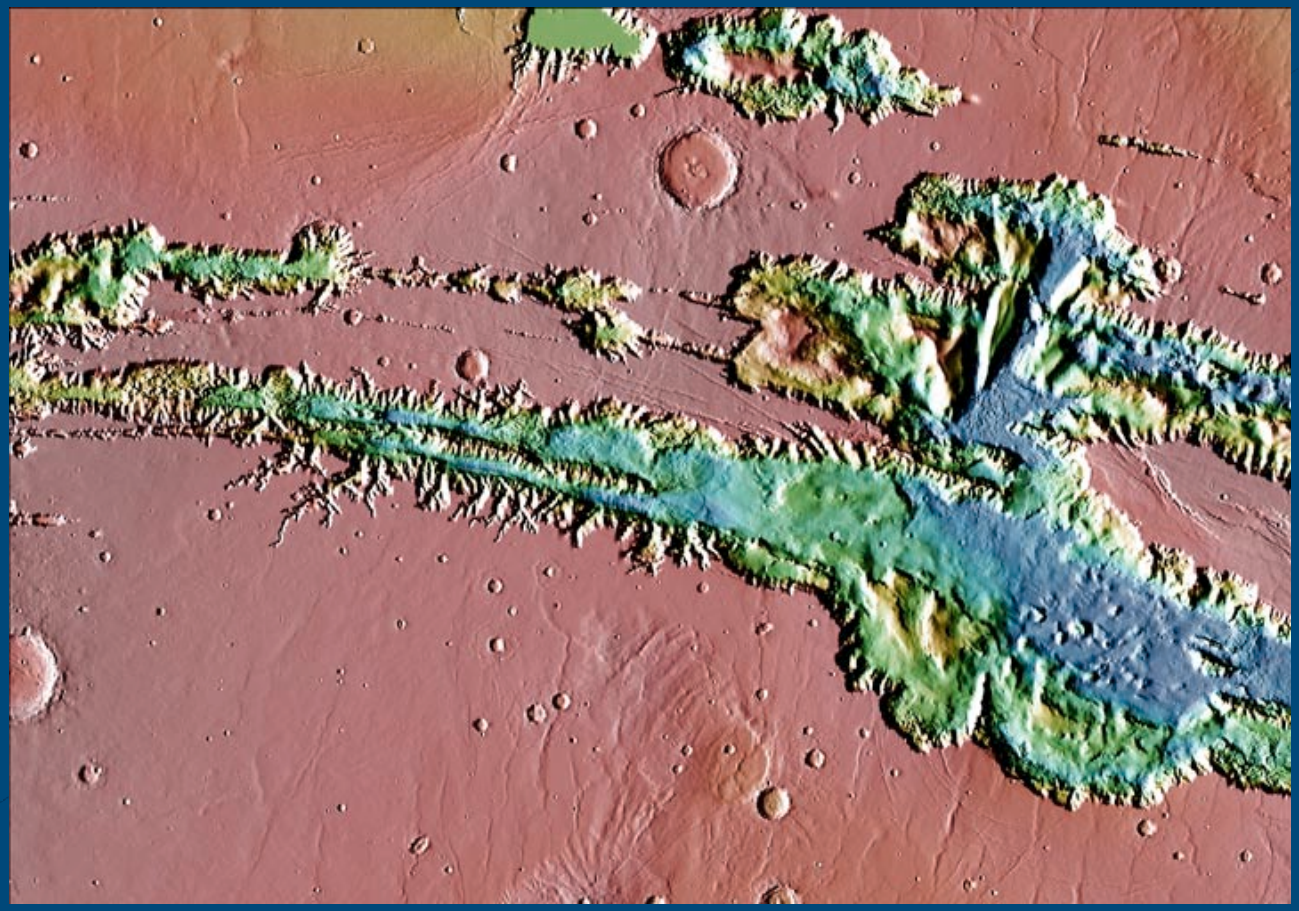
Mars Orbiter Laser Altimeter (MOLA) …The End of Photogrammetry?



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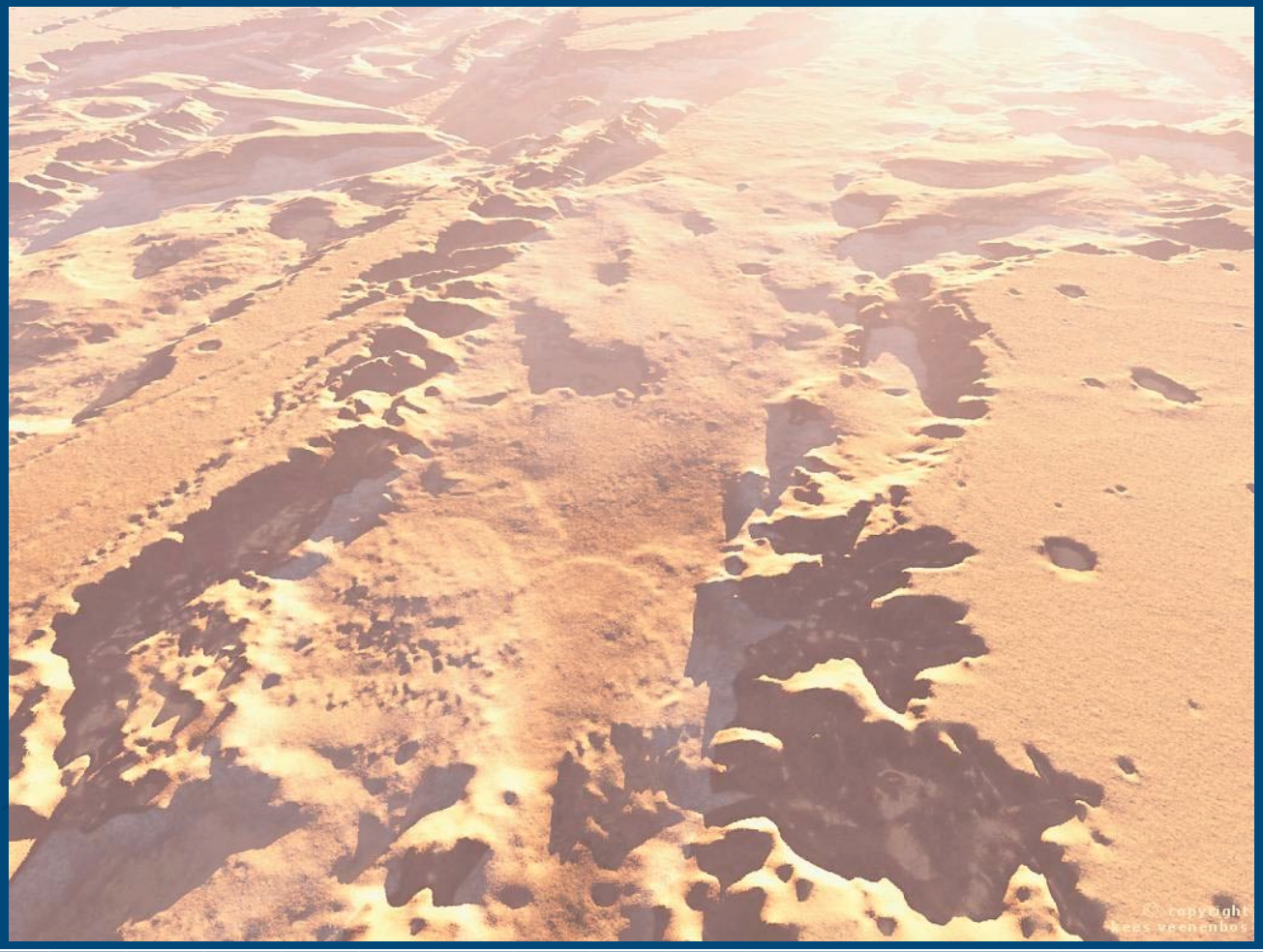
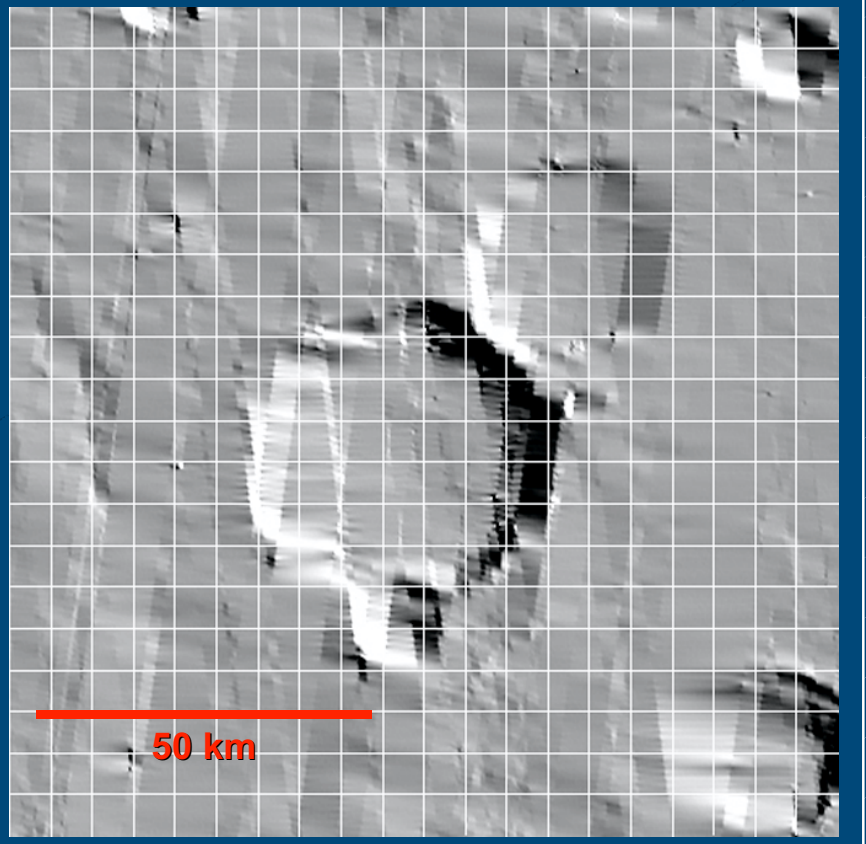


MOLA Statistics

600 Million Shots

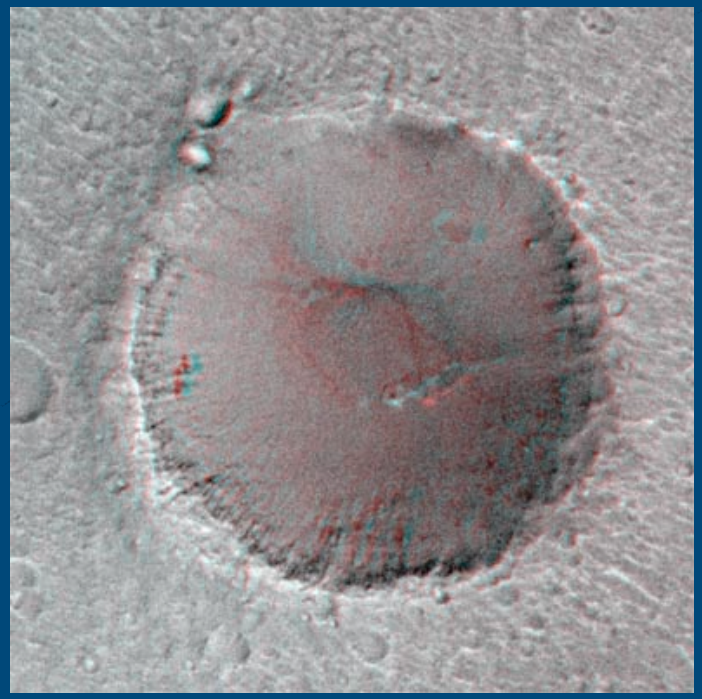
- <1 m vertical precision
- ~10 m vertical accuracy
- 160 m spot diameter
- 300 m along-track spacing
- 1–5 km gaps between tracks



MOC-NA Stereo

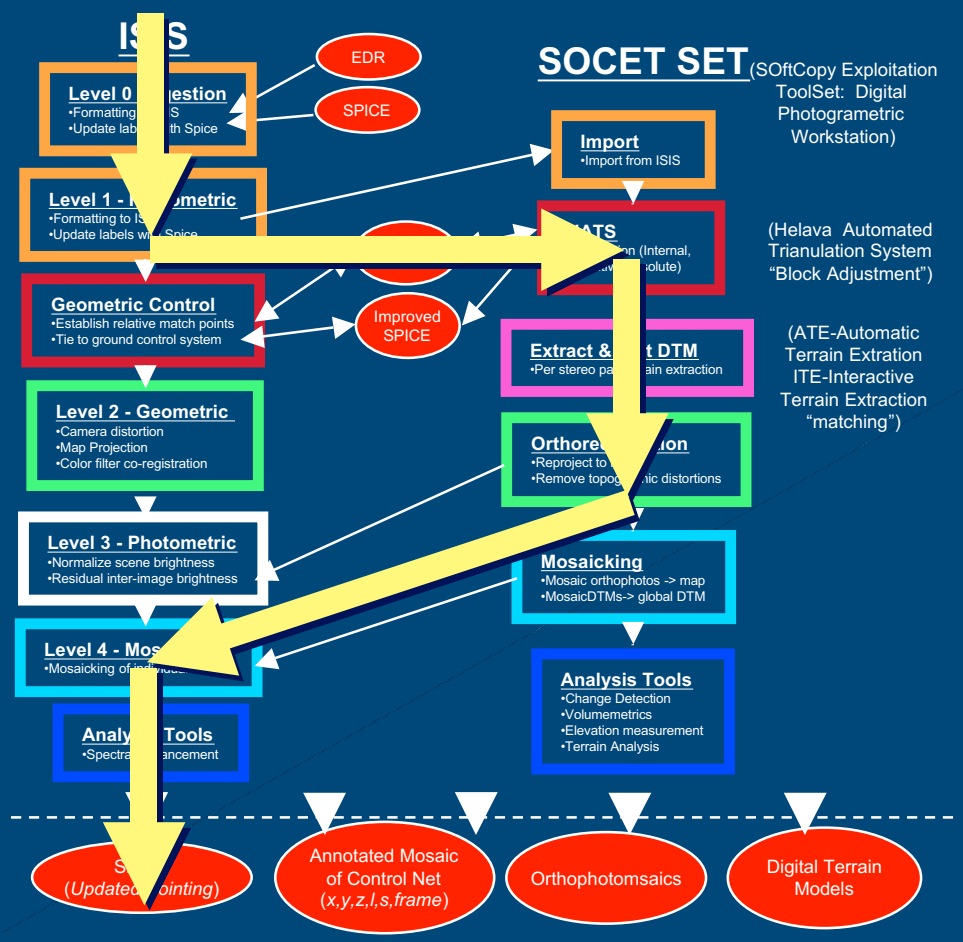
>150 Stereopairs

- Out of >30,000 images
- ~10^{-5} of Mars by area
- ≥1.5 m/pixel (3 m typical)
- DEM horizontal resolution ~3 pixels (10 m typical)
- 2 m vertical precision typical
  EP = 0.2 pixel / (base/ht)
- Accuracy depends on ties to MOLA



"Big Crater" (1.5 km diam)
Ares Vallis, S of Pathfinder site,
5 m/pixel

Stereo Process Flow
ISIS & SOCET SET



ISIS: Integrated Software for Imagers & Spectrometers

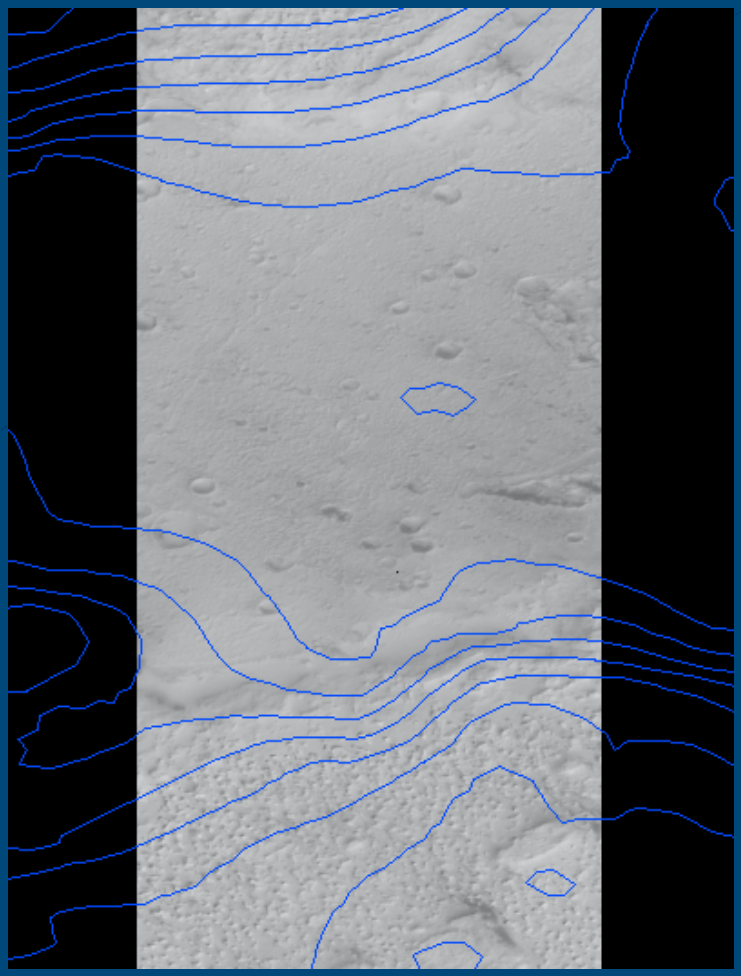
- USGS in house s/w
- Strengths:
  - Ingestion/calibration
  - Planimetric mosaicking
  - Quantitative analysis including spectroscopy
- Use for calibration, 2D proc.

SOCET SET: SOftCopy Exploitation ToolSET

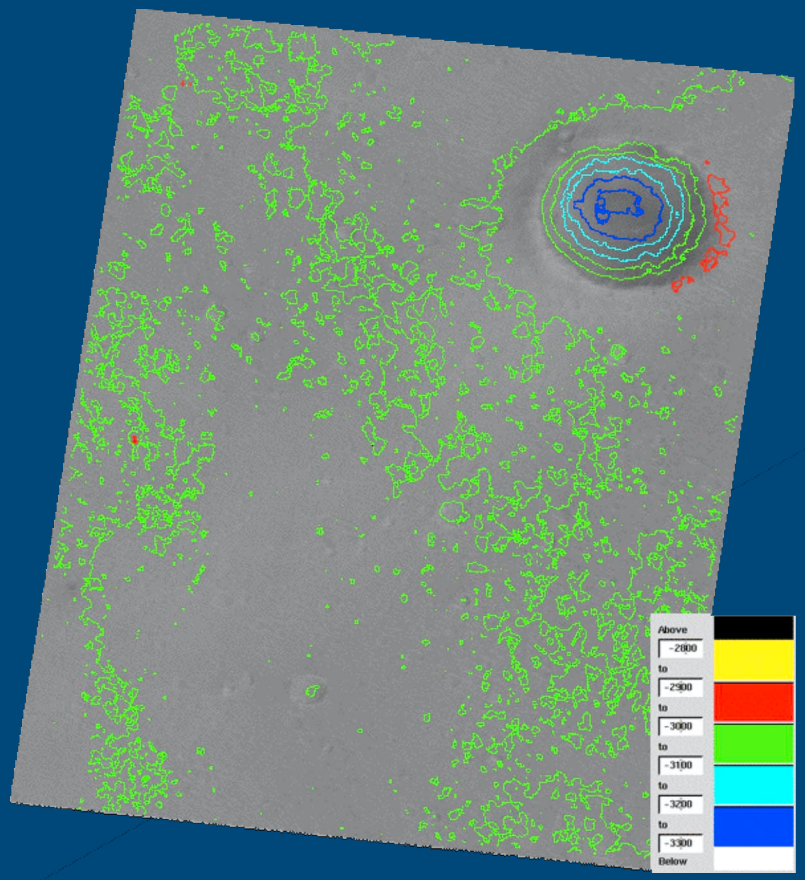
- LH Systems commercial s/w
- Strengths:
  - Stereo display/input
  - Bundle-block adjustment
  - Automatic DTM matching
- Use for 3D processing
- Write new sensor models

MOC Stereomapping Approach

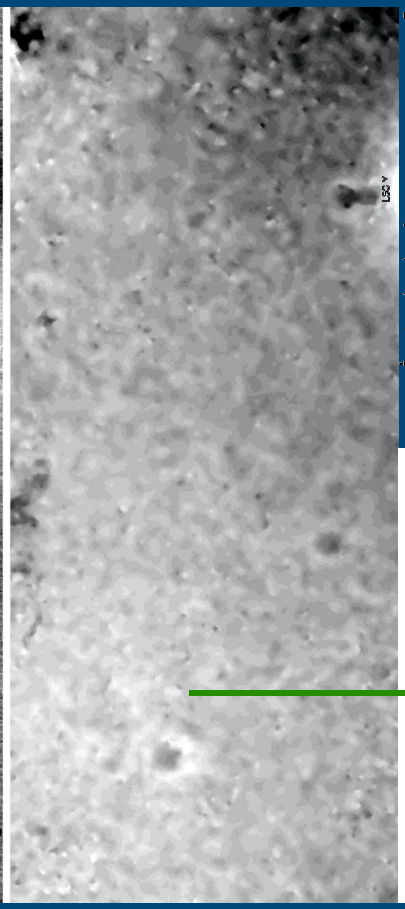
- Geometric calibration of NA, alignment of NA, WA to MOLA using preflight, inflight data
- Decompression, QC, and radiometric calibration with ISIS
- ISIS program to translate metadata from NAIF SPICE Kernels to SOCET support file
- SOCET “Generic Pushbroom” sensor model (physically rigorous but simple)
- Control images to high resolution MOLA DEM produced in ArcView
- Collect & edit DEM as usual
- Analyze w ArcView, custom s/w

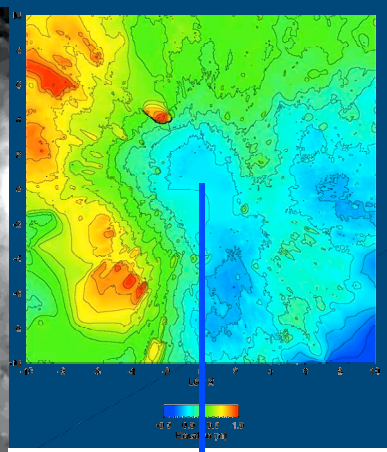


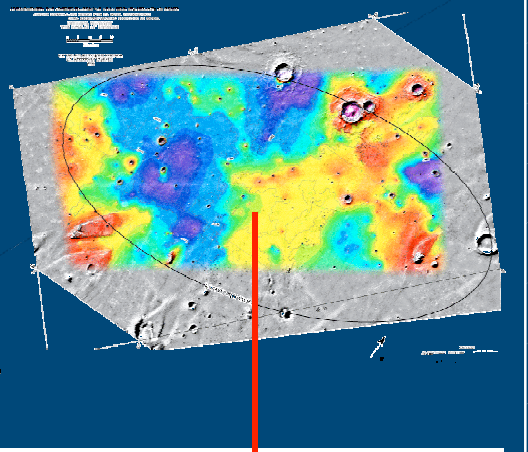
First MOC DEM: “Big Crater”

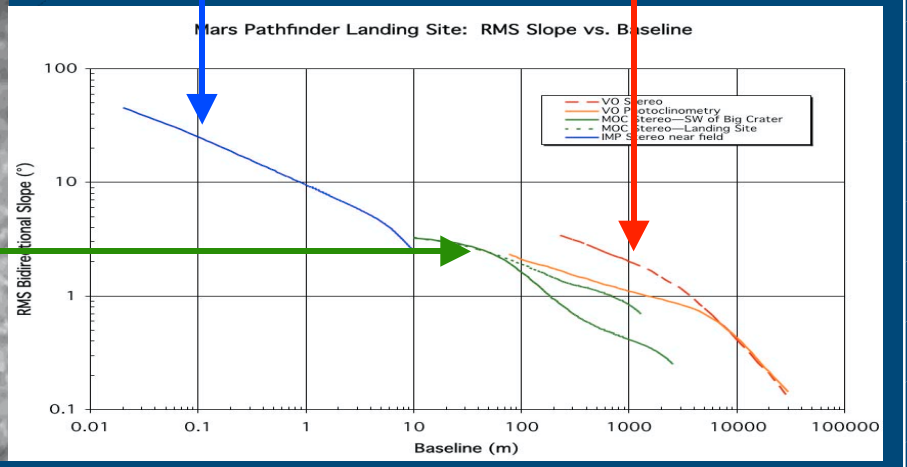


Pathfinder Site: "Ground Truth" Comparison of Slope v. Baseline

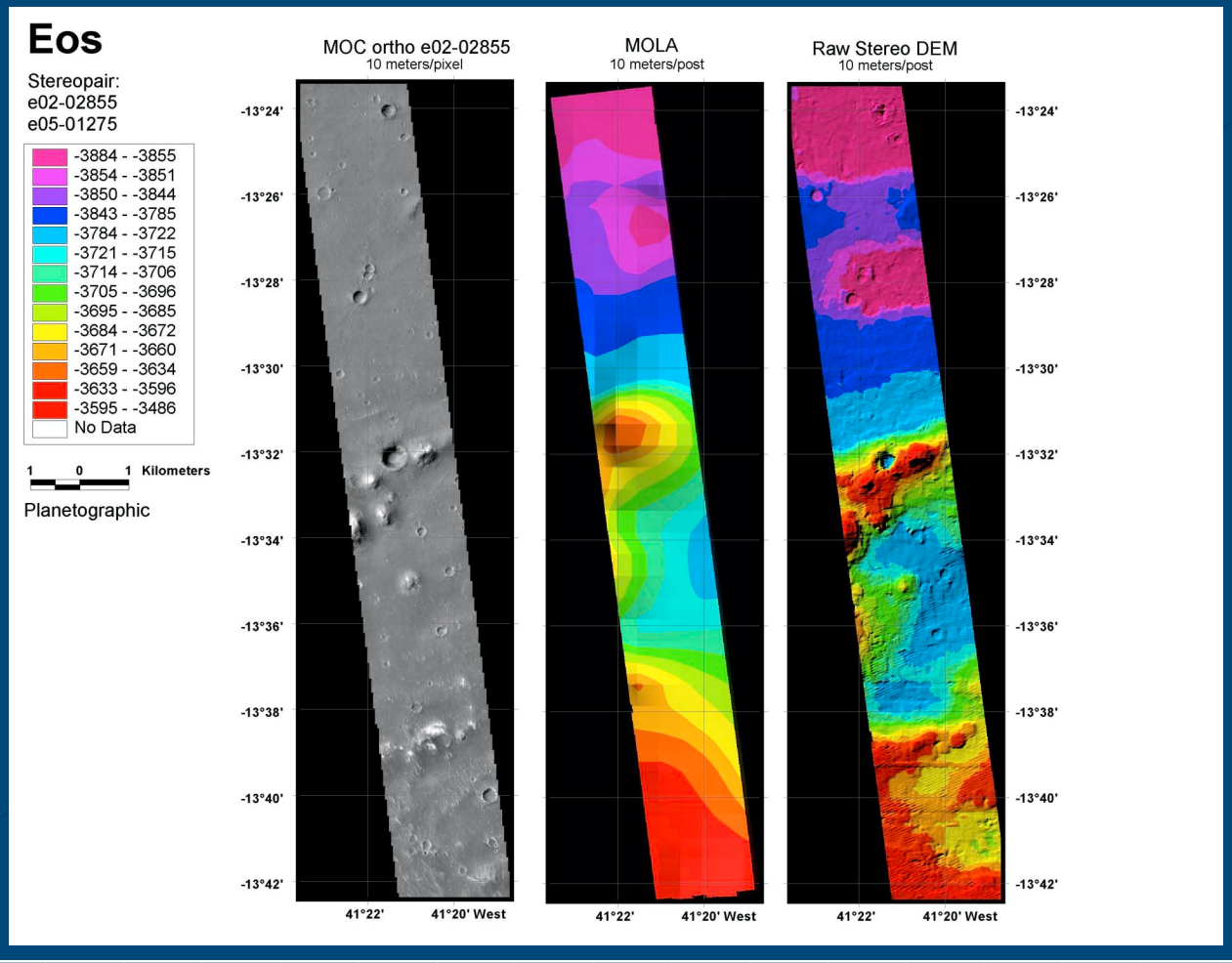




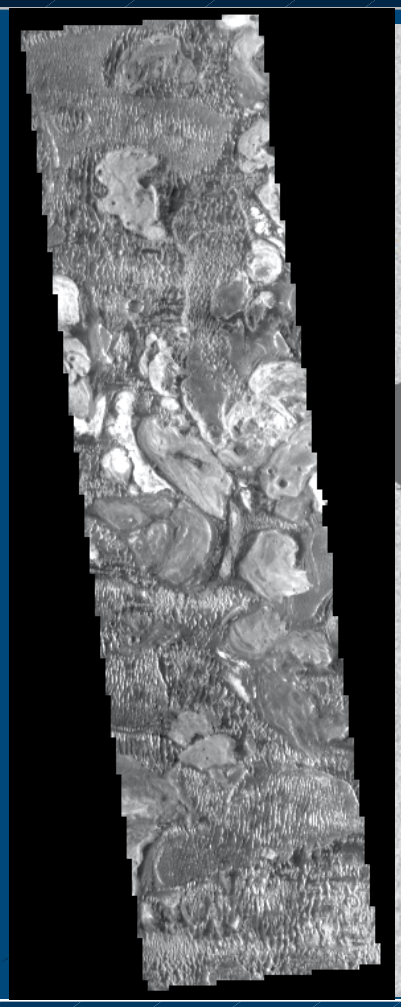


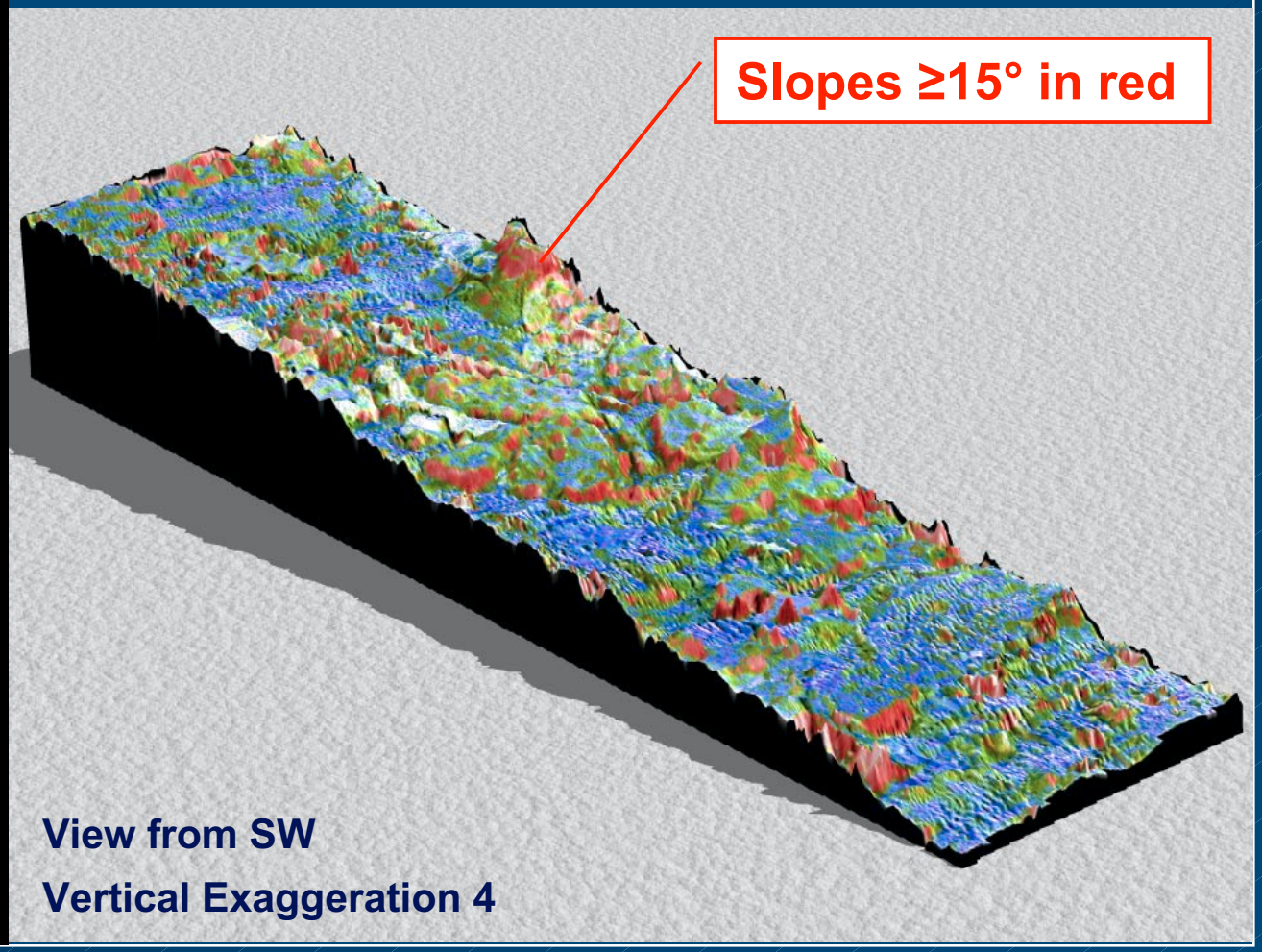


Eos Chasma: E02-02855/E04-01275



Melas Chasma Visualized



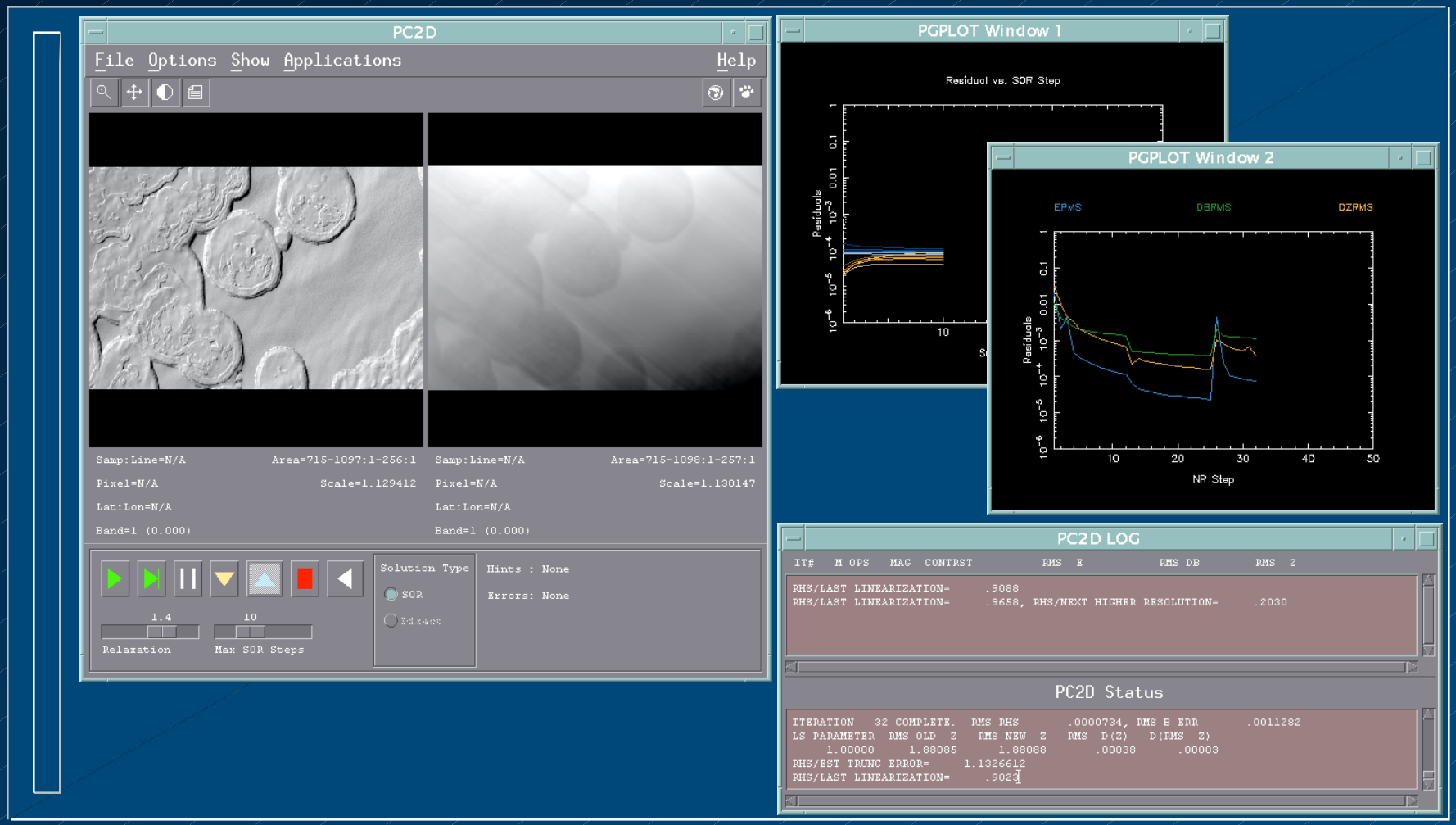


Photoclinometry: Shape from Shading

Inference of topo from brightness information in image

- Highly complementary to stereo
- Brightness <–> slope; must integrate to get heights
  - Most accurate for short baselines, down to 1 pixel
- Can be formulated various ways
  - 1 D (profile): must make assumptions about slope direction
  - 2 D (DEM): use implicit or explicit continuity of surface
    - Think of it as a nonlinear PDE with image as driver
    - Think of it as the inverse problem to hillshading
    - Think of it as nonlinear lsq fitting of model image to real image
- 2D solution is iterative, delicate, CPU+labor intensive
- Results depend on validity of photometric model used
  - Surface reflectivity variations –> artifacts
  - Surface photometric function
  - Atmospheric scattering: amount and directionality

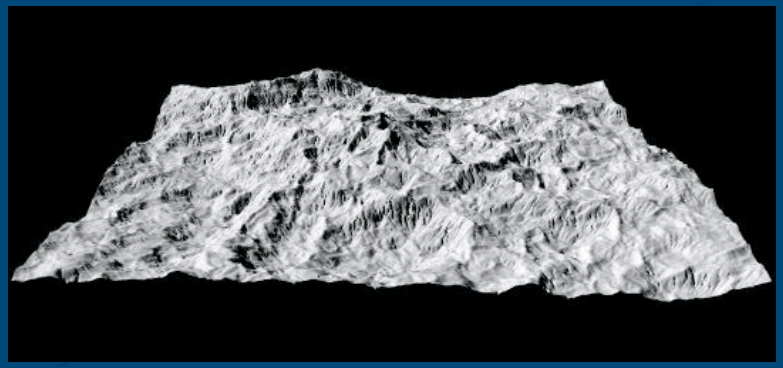
2 D Photoclinometry
ISIS Graphical User Interface



Effect of Haze and Albedo

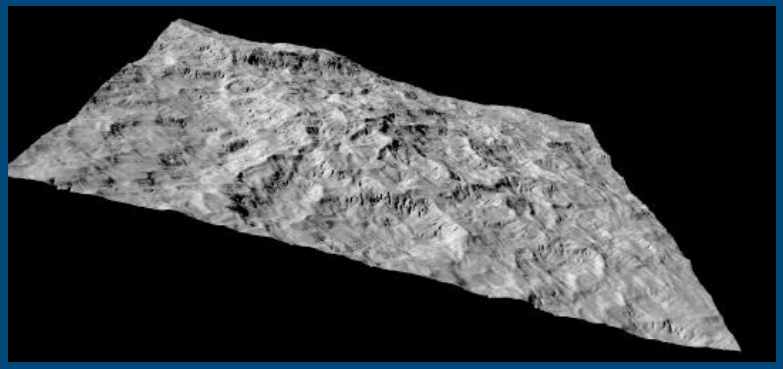
Sun is from upper left in all examples

Correct Haze and Albedo



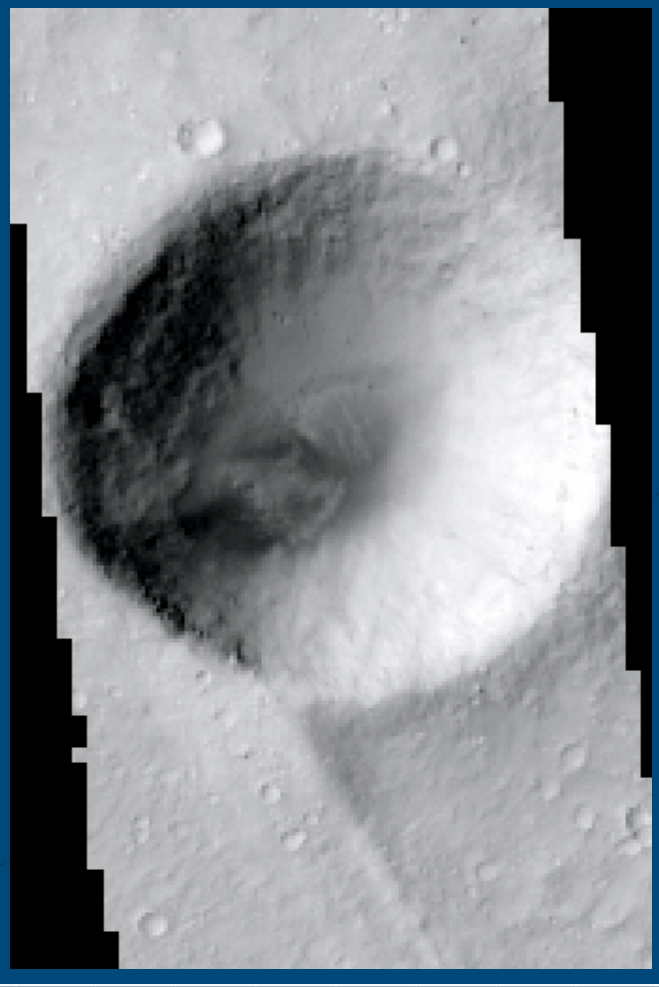
Too much Haze subtracted

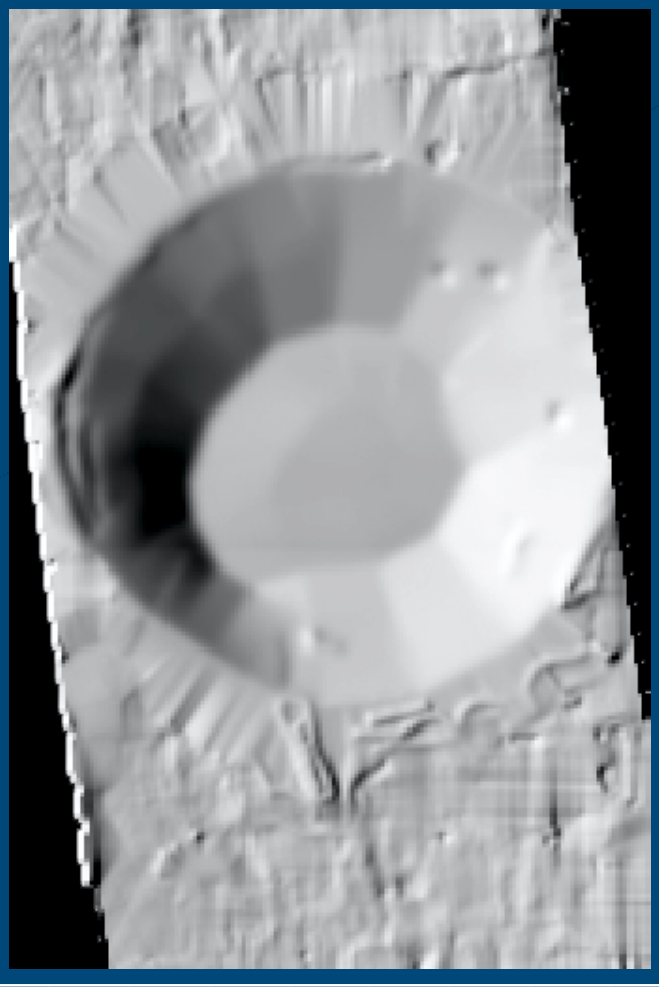
Albedo underestimated



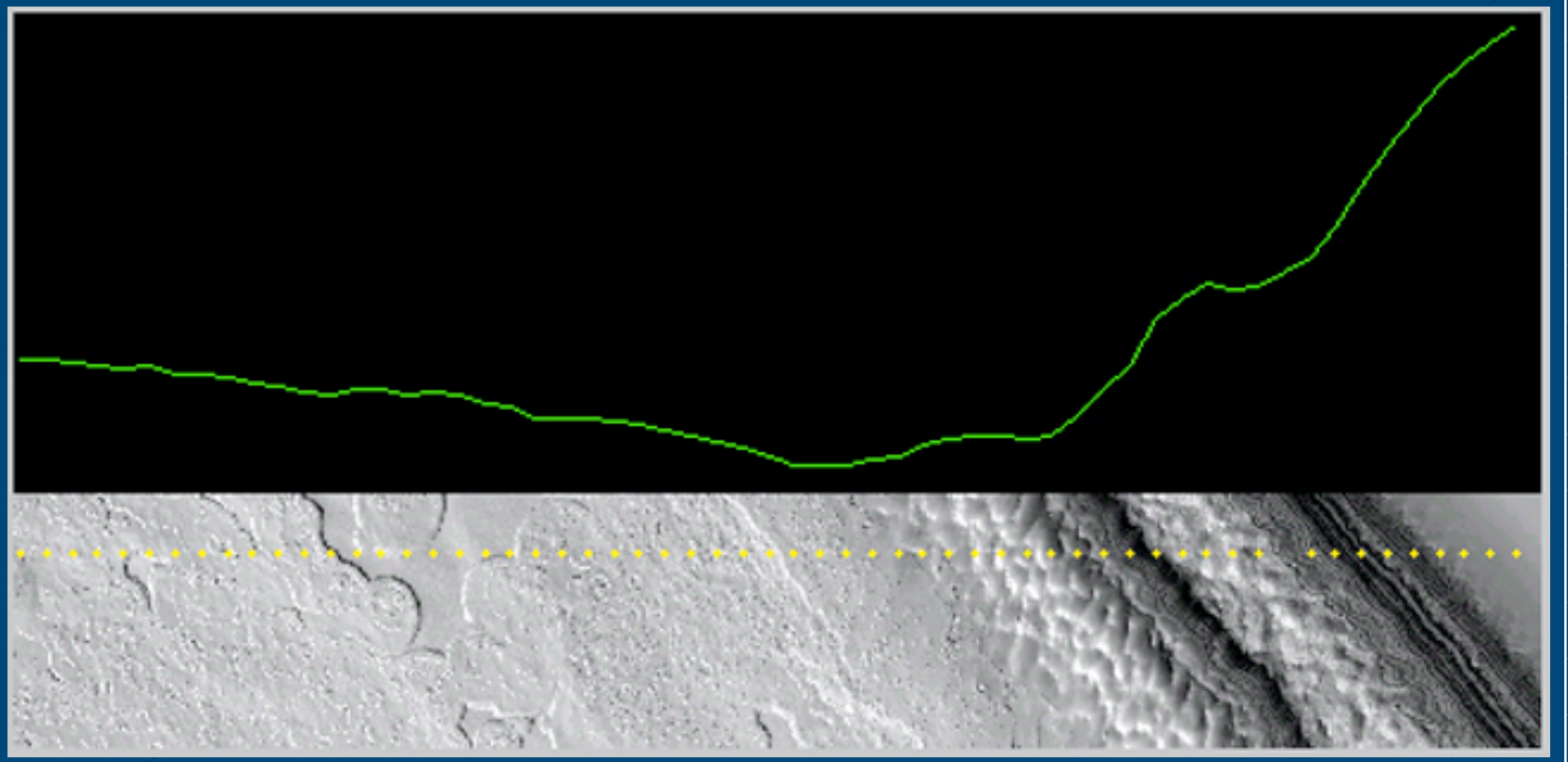
Albedo overestimated

Forward Fitting for Haze: Simulate Image from DEM





Inverse Fitting for Haze: Tweak Photoclinometry to Match DEM



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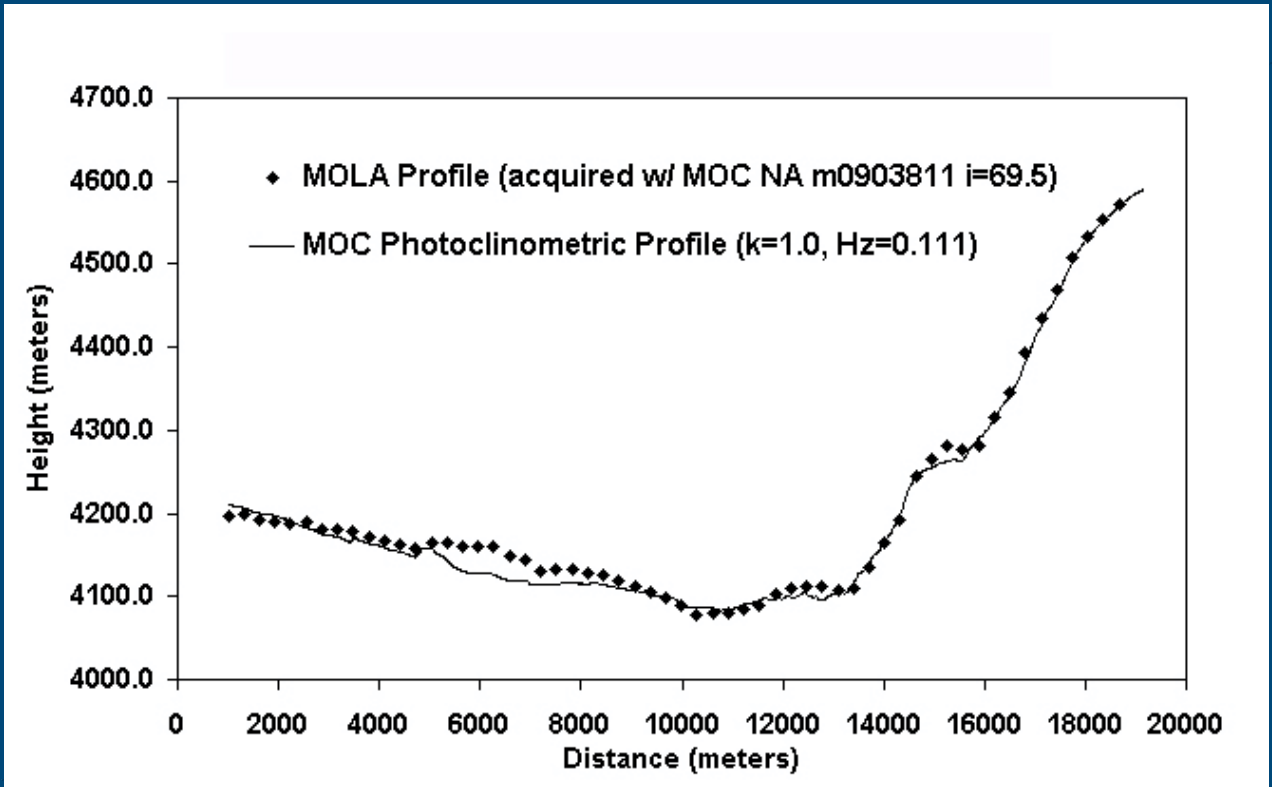

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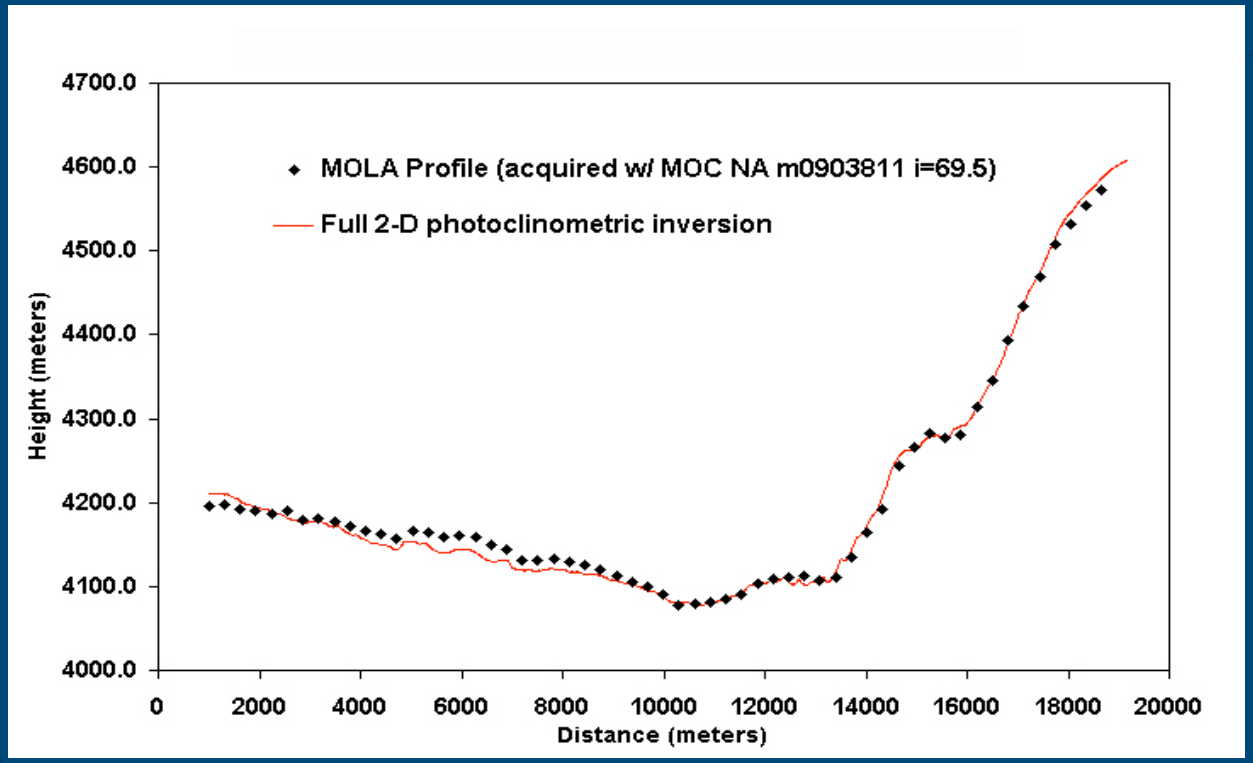
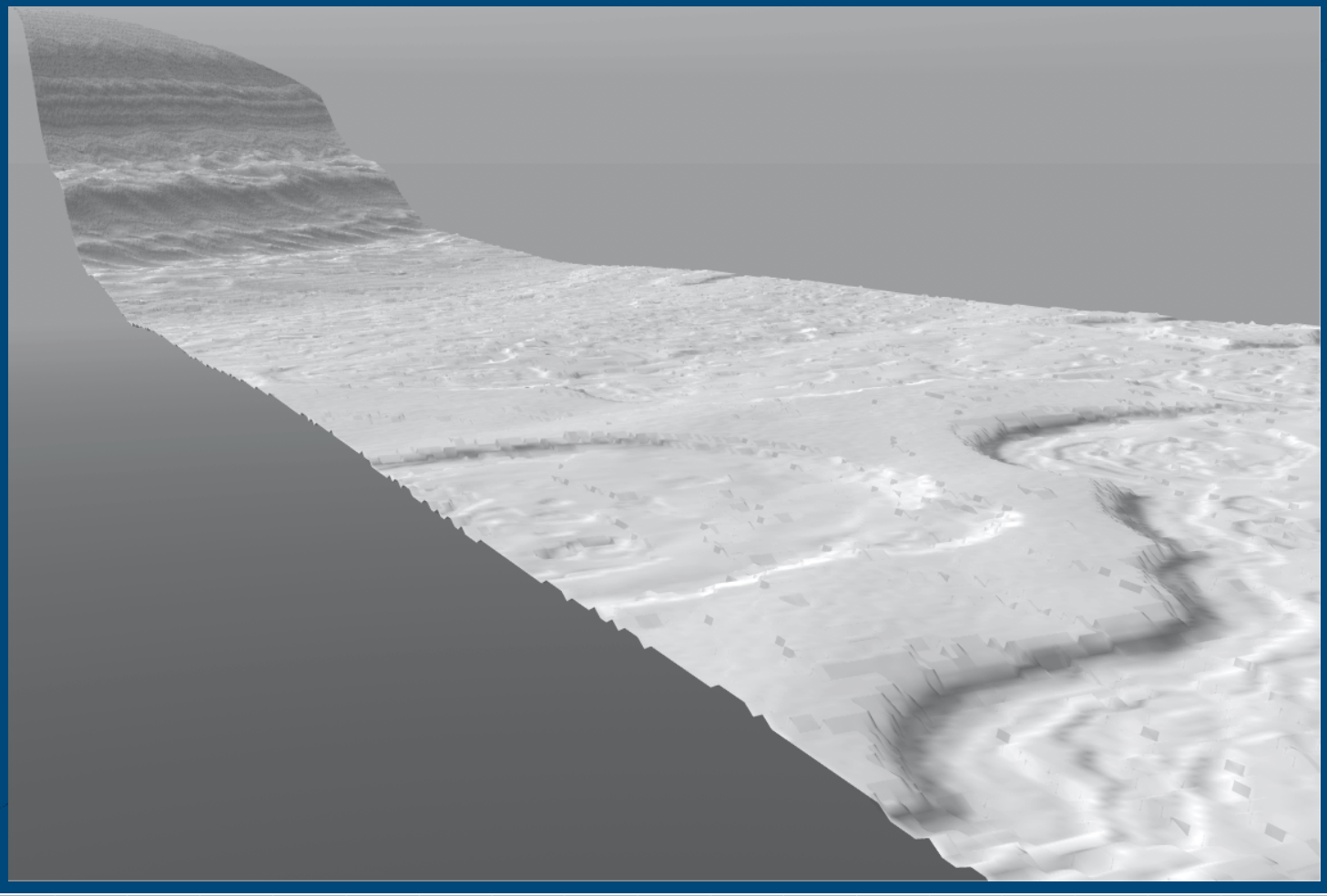


Perspective Views of Polar Layered Terrain



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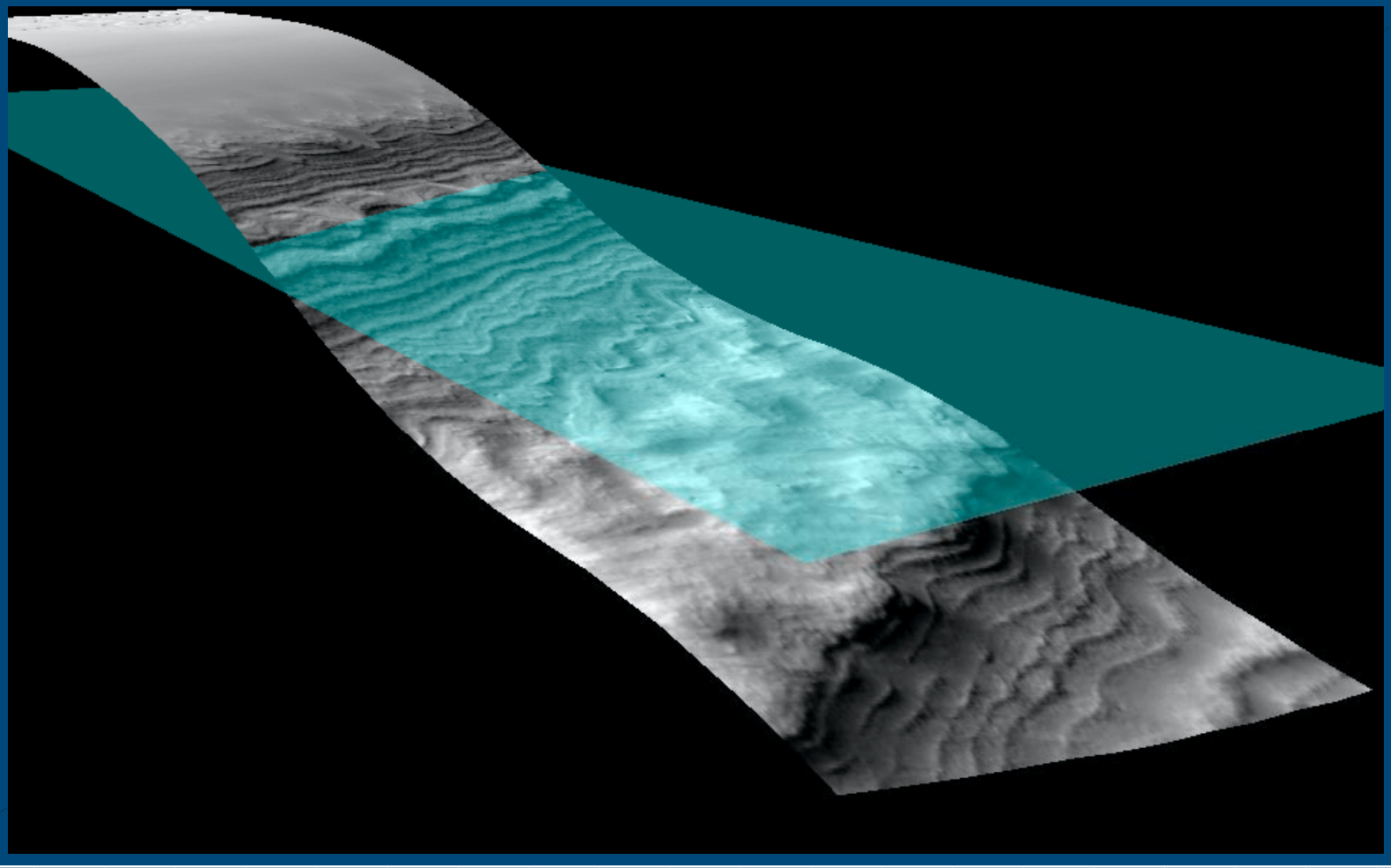

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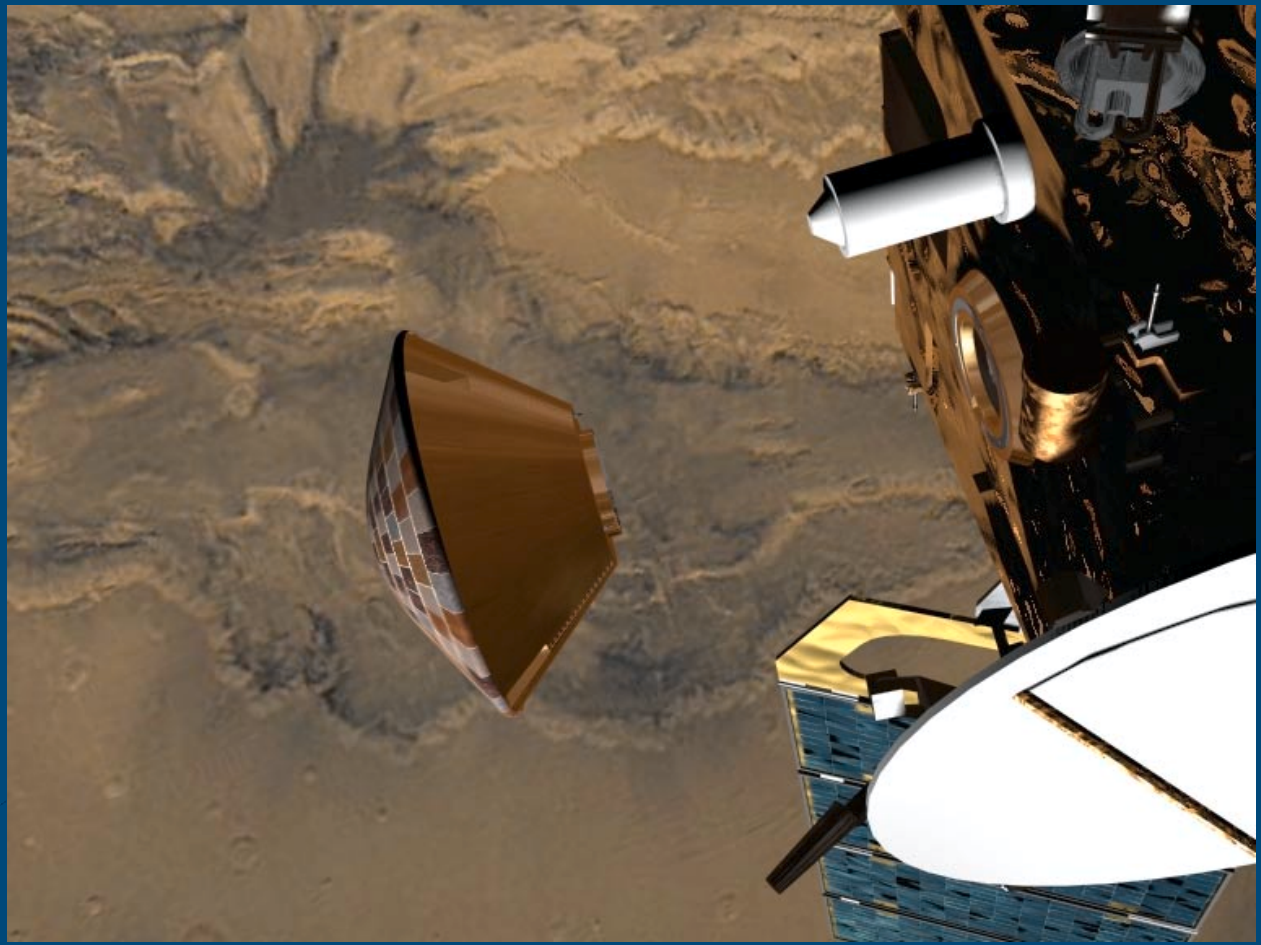
Mars Express HRSC (2003)



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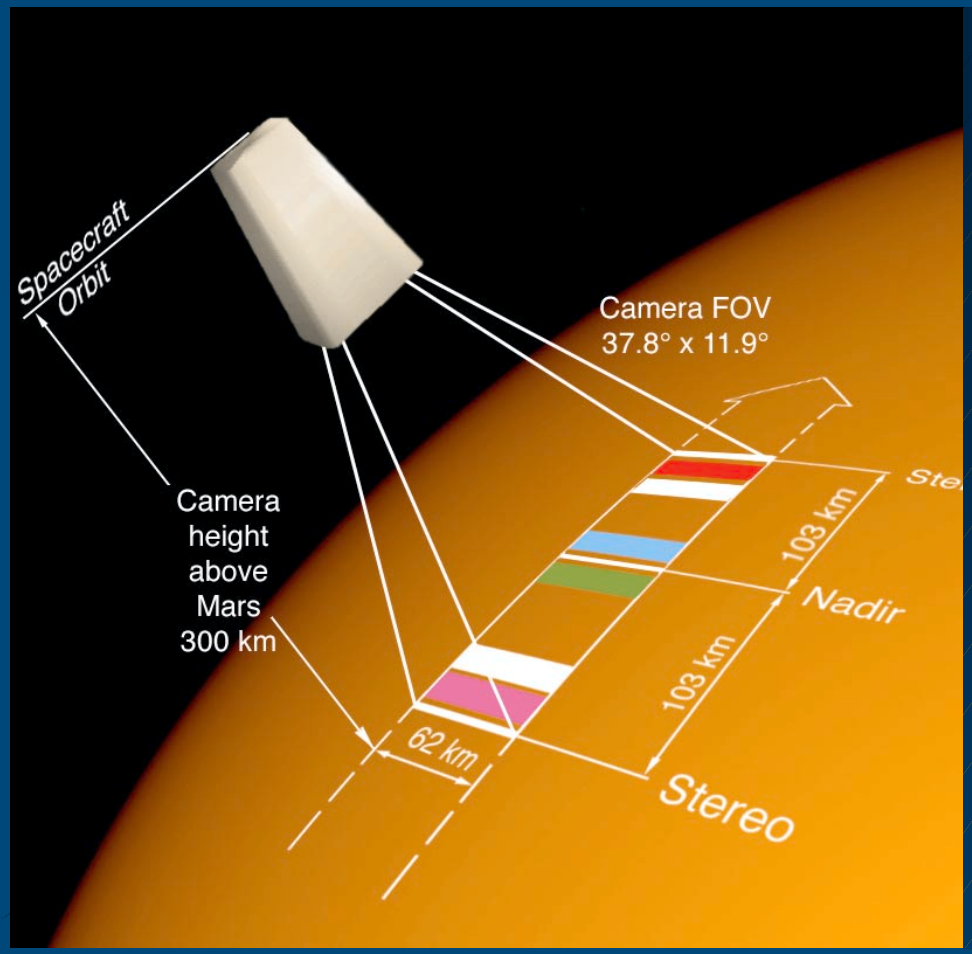
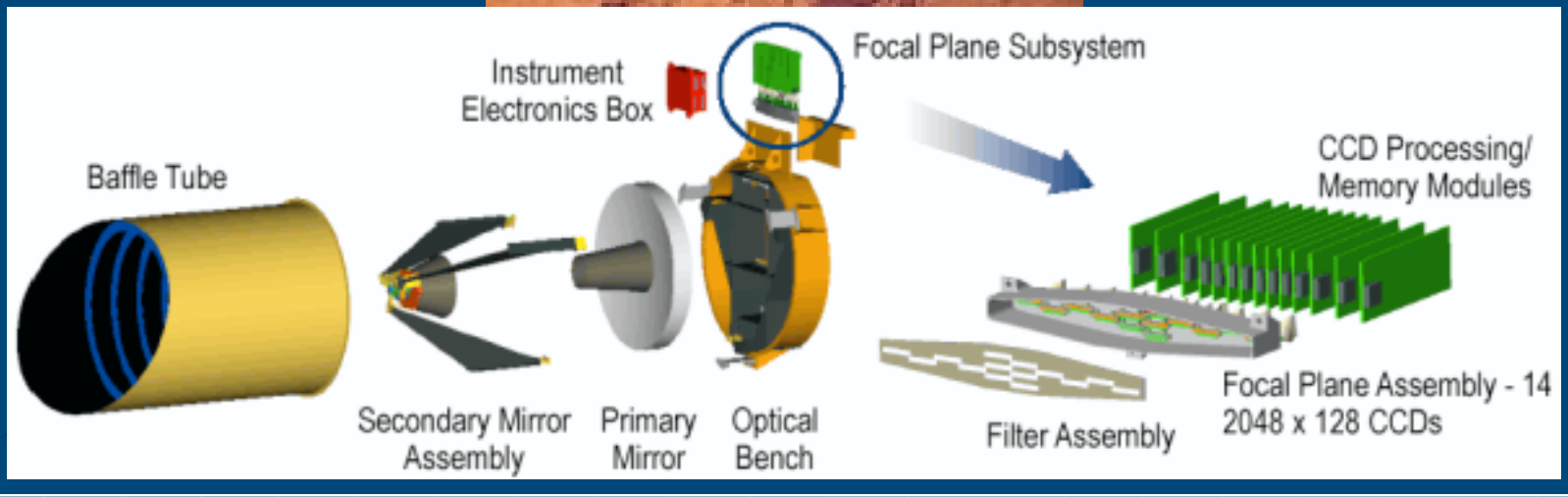


First Dedicated Stereo Camera at Mars

- 9 pushbroom lines
- 3 stereo lines+color
- 5000 pixel swath
- 10 m/pixel at periapse
- *Many* % of Mars with
  - ≤50 m DEM spacing
  - ≤10 m vertical prec
- Plus larger areas at poorer resolutions

MRO—HiRISE (2005)



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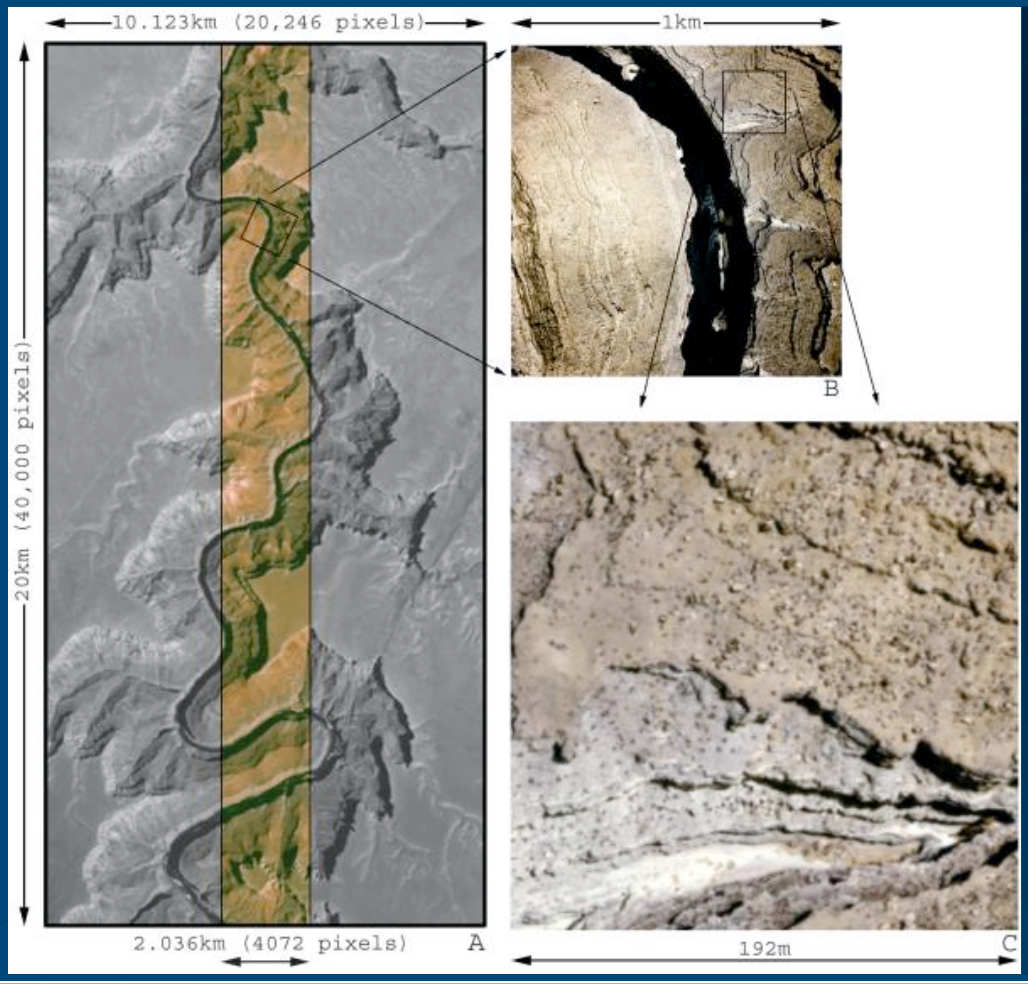
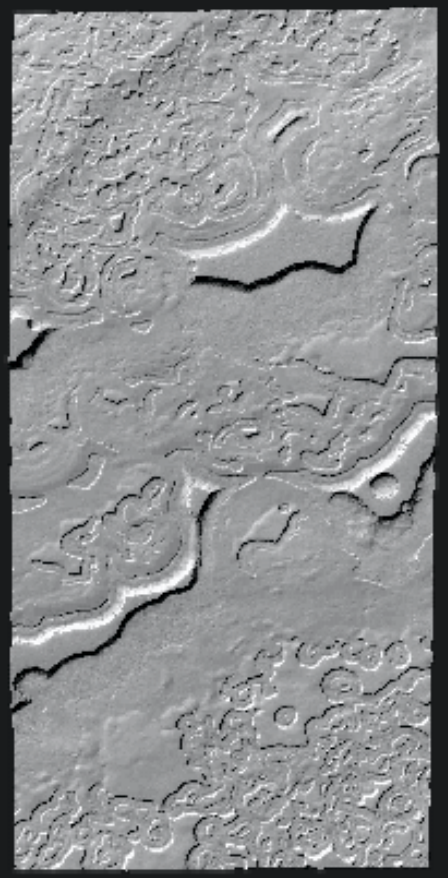
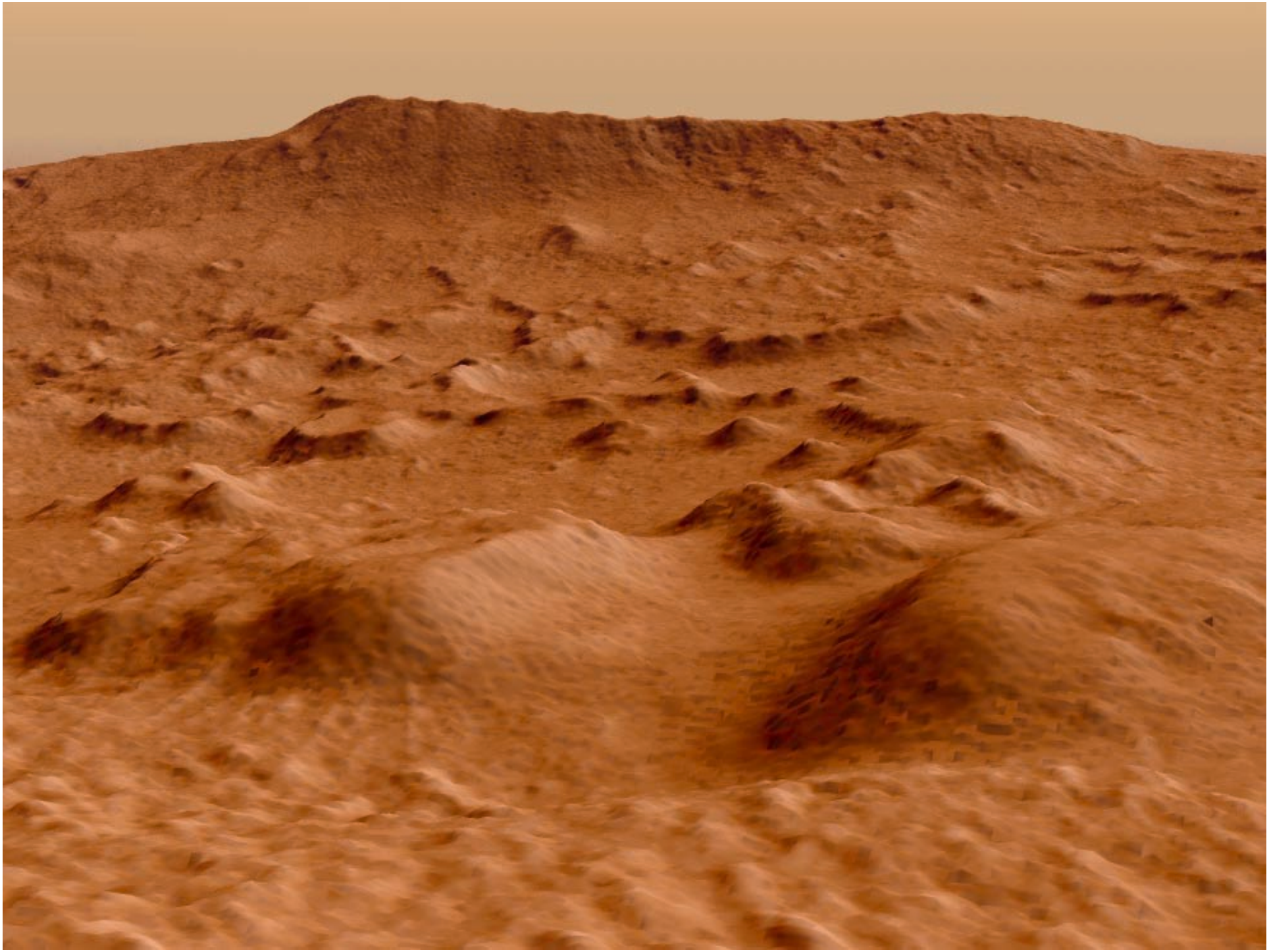


Ultrahigh Resolution

- 20,000 pixel swath
- 4000 pixels in color
- 30 cm/pixel
- 1 m DEM spacing
- 20 cm vertical prec
- Stereo a priority: ~1000 pairs
- ~10^{-3} of Mars in stereo

“Flyover” of Polar Layered Terrain